

COSMIC RAY MUONS

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Outline

- ❖ What Is a Muon?
- ❖ Cosmic Origin and Interactions
- ❖ Muon Production and Path
- ❖ Effects in The Atmosphere
- ❖ Using Muons to Test Beam Profile Monitor
- ❖ Conclusion

What Is a Muon?

- ❖ Discovered in 1936
 - ❖ Carl D. Anderson and Seth Neddermeyer

- ❖ Fundamental Particle

- ❖ Similar to Electron
 - ❖ Lepton
 - ❖ Mass is ~ 207 Times That of the Electron

Interaction and Decay Modes

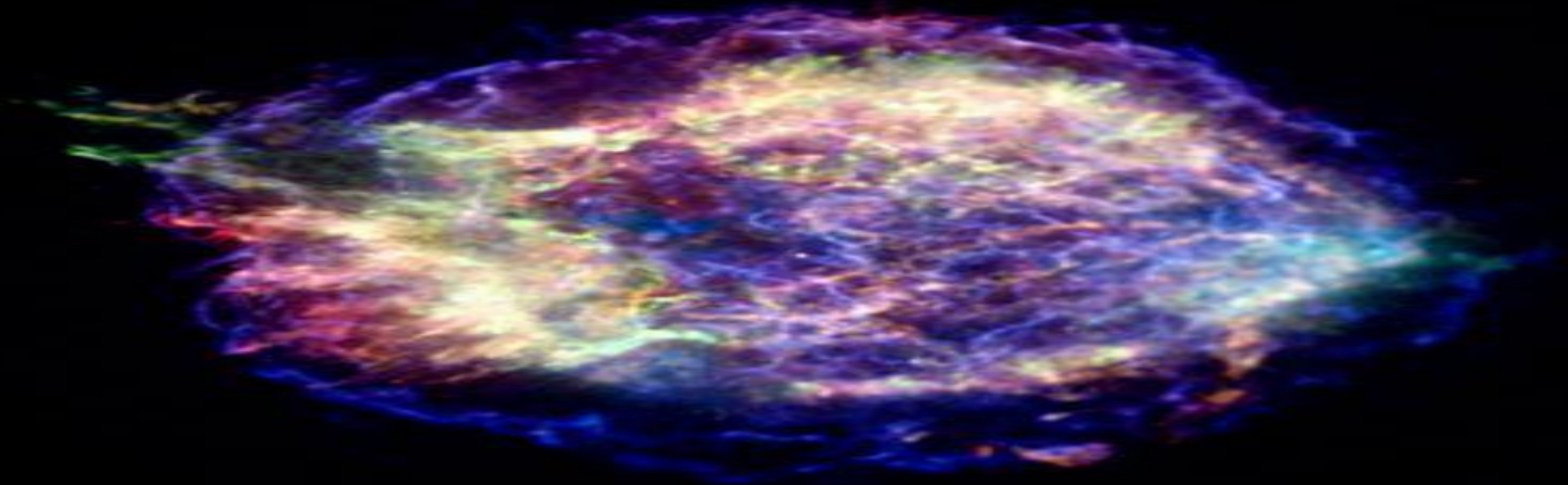
- ❖ Interaction
 - ❖ Gravitation
 - ❖ Electromagnetism
 - ❖ Weak Interaction
- ❖ Primary Decay Mode
 - ❖ $\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu$
 - ❖ $\sim 100\%$

Fundamental Properties

Property	Value
Muon Mass	$105.6583668 \pm 0.0000038$ MeV
Muon Electric Charge	e^- , e^+ (anti-muon)
Mean Life	2.19703 ± 0.00004 μ seconds
Spin	1/2
Magnetic Moment Ratio, μ/p	$3.18334539 \pm 0.00000010$
Electric Dipole Moment	$3.7 \pm 3.4 (10^{-19} \text{ecm})$

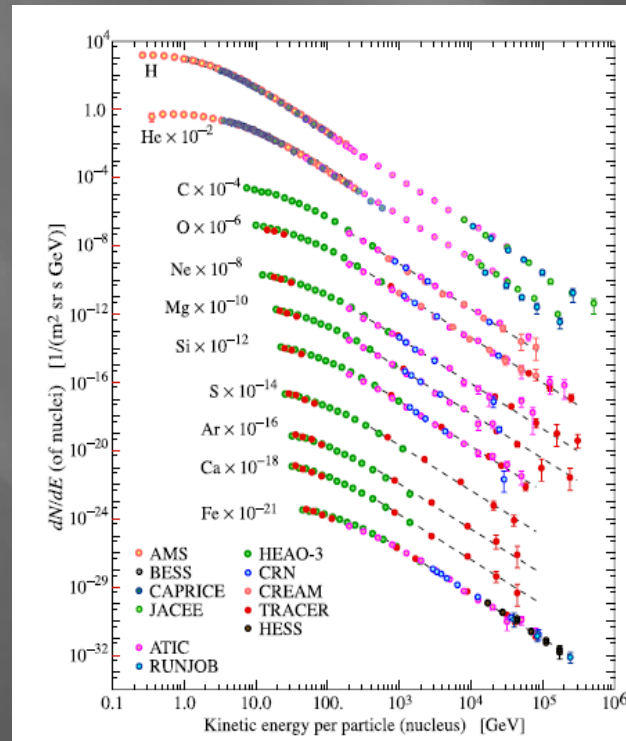
Cosmic Origin and Interactions

- ❖ Cosmic ray muons are created when high energy primary cosmic rays interact with earths atmosphere.
- ❖ The primary sources of cosmic rays are supernovae



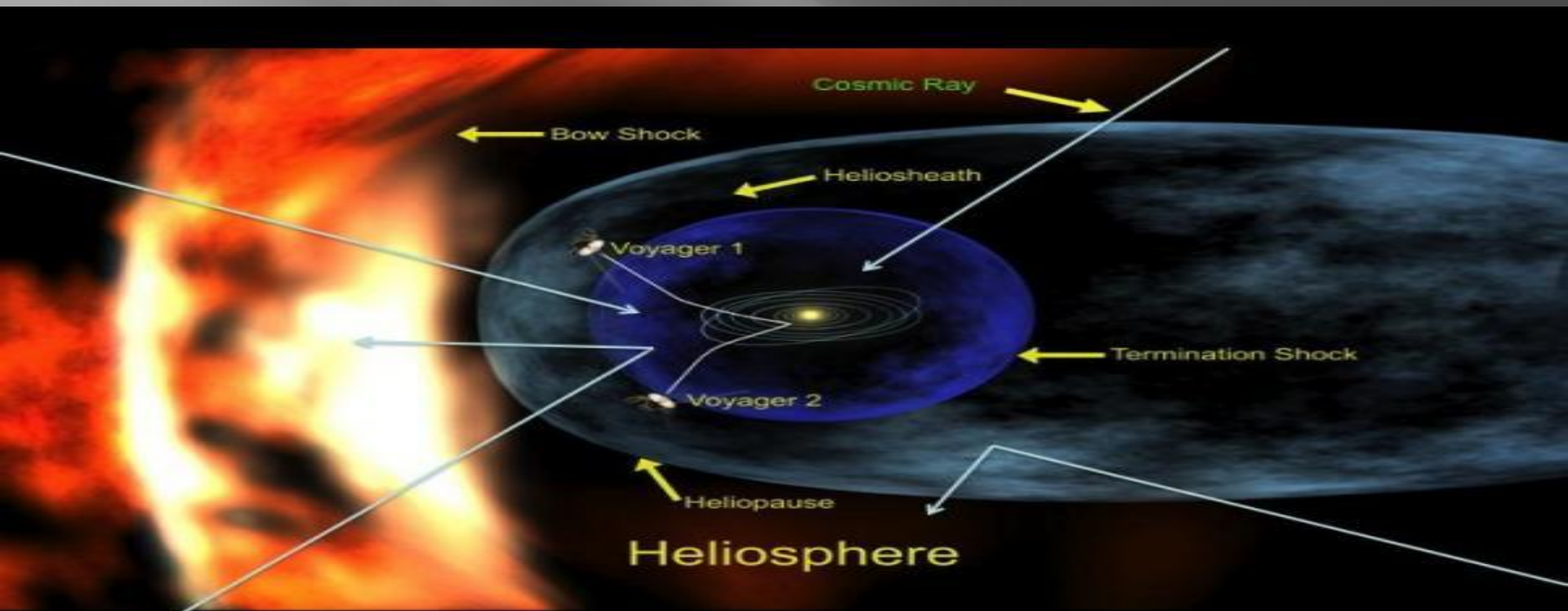
Primary Cosmic Rays

- ❖ Components
 - ❖ 89% Hydrogen Nuclei (Protons)
 - ❖ Remaining 11% includes Helium, Carbon and Oxygen among other less abundant elements



Primary Cosmic Ray Interactions

- ❖ Gravitation
- ❖ Electromagnetism



Lorentz Force Law

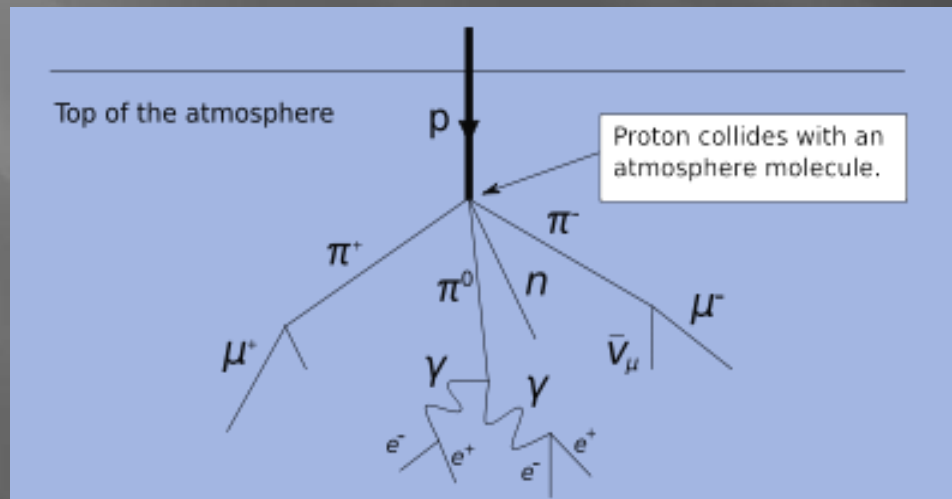
- ❖ The equations of motion for primary cosmic rays interacting with earth's magnetic field can be derived from the Lorentz force law.

$$F = ma = q(v \times B)$$

- ❖ Manuel Vallarta
 - ❖ Earth is approximated as a magnetic dipole

Production and Path

- ❖ High energy primary cosmic rays collide with atmospheric molecules at an altitude of $\sim 15\text{km}$ to produce secondary cosmic rays
- ❖ Most abundant collisions
 - ❖ Proton $\rightarrow N_2$ Diatomic Molecule
 - ❖ Proton $\rightarrow O_2$ Diatomic Molecule



Components of Secondary Cosmic Rays

- ❖ Protons, Neutrons, Pions, Kaons, Muons, Electrons and Photons
- ❖ Muons are the decay product of Pions and Kaons
 - ❖ $\pi^{-} \rightarrow \mu^{-} + \bar{\nu}_{\mu}$
 - ❖ $K^0 \rightarrow \pi^{+} + \pi^{-}$

Path Through The Atmosphere

- ❖ The mean energy of muons at the site of production (~15 km alt.) is 6 GeV. This energy corresponds to a velocity of .9998c. This velocity is derived from the relativistic equation for energy

$$E_{\mu}(v) = \frac{m_{\mu_0} c^2}{\sqrt{1 - \frac{v_{\mu}^2}{c^2}}}$$

Path Through The Atmosphere

- ❖ Time Dilation

- ❖ The lifetime of a 6 GeV cosmic muon as measured from Earth frame

$$\tau = \gamma \tau_0 = 109.856 \times 10^{-6} \text{ s}$$

- ❖ The incident angle of the majority of muons at sea level is close to the zero zenith angle

- ❖ $\sim 1 \text{ muon / cm}^2 \text{ minute}$

Effects in the Atmosphere

- ❖ Temperature Variance
- ❖ Interaction With Thunderstorms

Effects Cont'd

- ❖ Temperature Variance
 - ❖ Cosmic muon flux intensity corresponds directly with the atmospheric temperature at different pressures, altitudes and zenith angles

Effects Cont'd

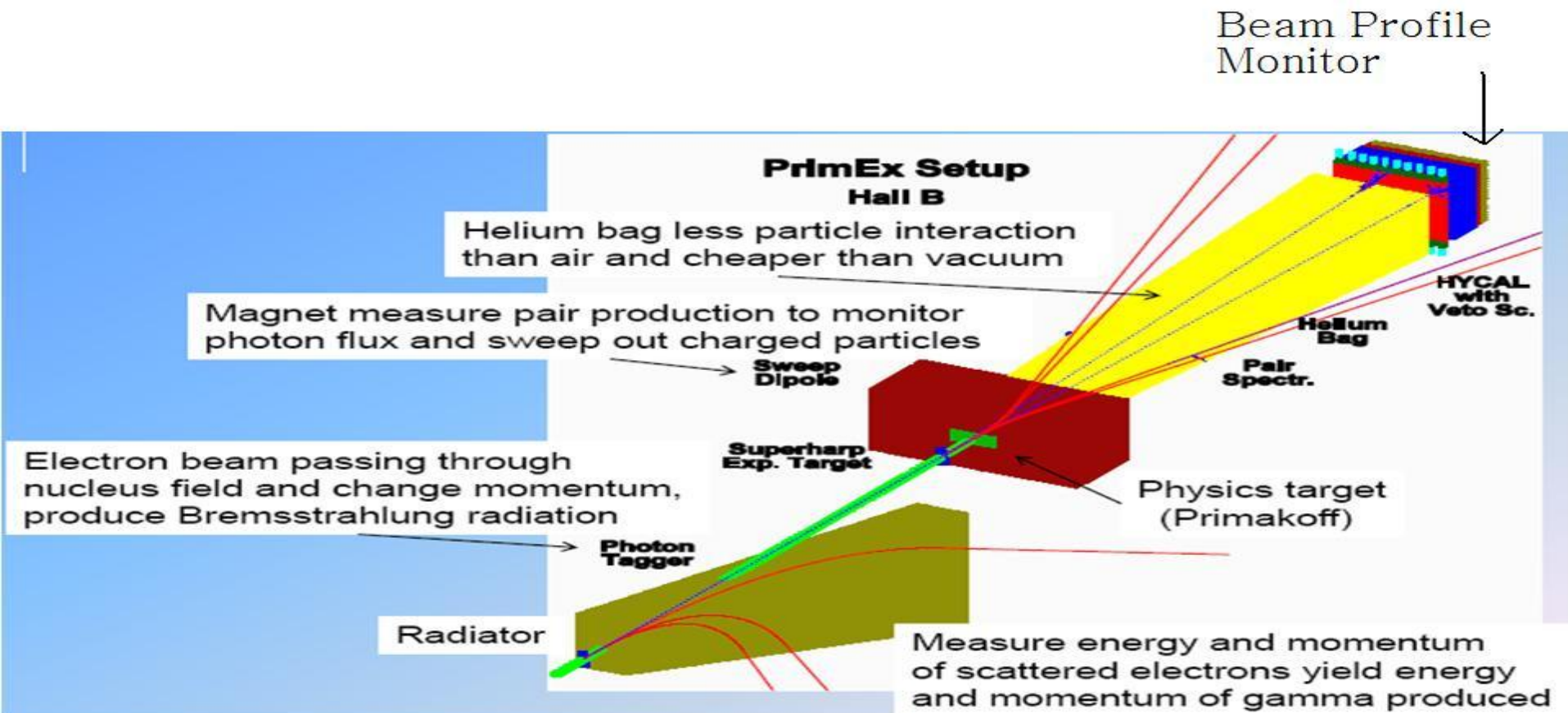
- ❖ Interaction With Thunderstorms
 - ❖ The mean muon flux intensity has been observed to decrease during and near thunderstorms
 - ❖ Lightning also effects the mean muon flux intensity during these thunderstorms

Energy Loss In The Atmosphere

- ❖ Ionization and excitation processes
- ❖ Brehmsstrahlung
- ❖ Photonuclear and photonucleon interactions
- ❖ Primary mode of energy loss for low energy muons (~ 6 GeV) is through Ionization and Brehmsstrahlung

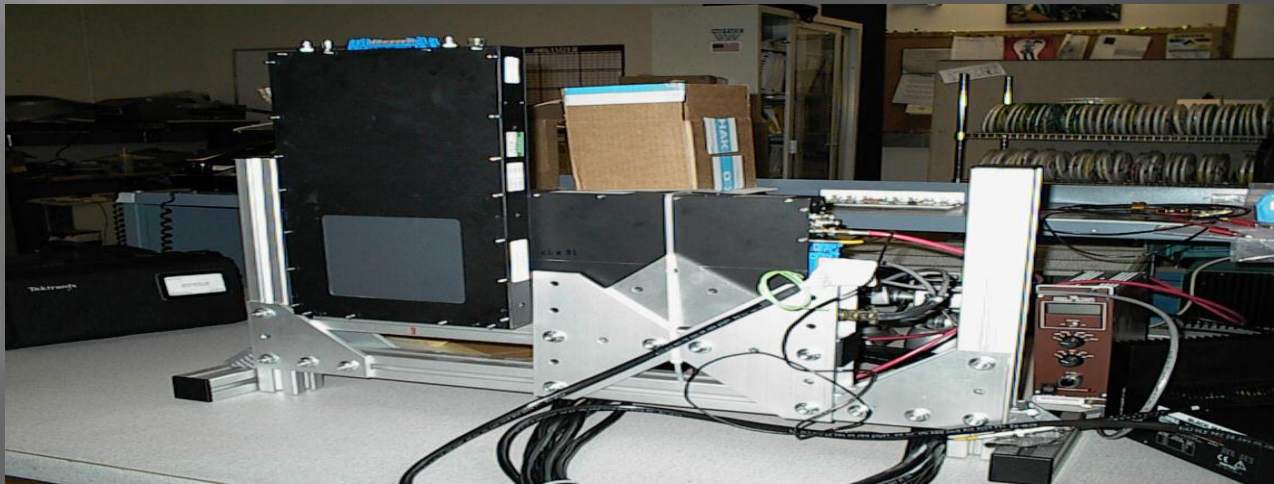
Testing the Beam Profile Monitor

- ❖ Thomas Jefferson National Laboratory
- ❖ PrimEx Experimental Setup



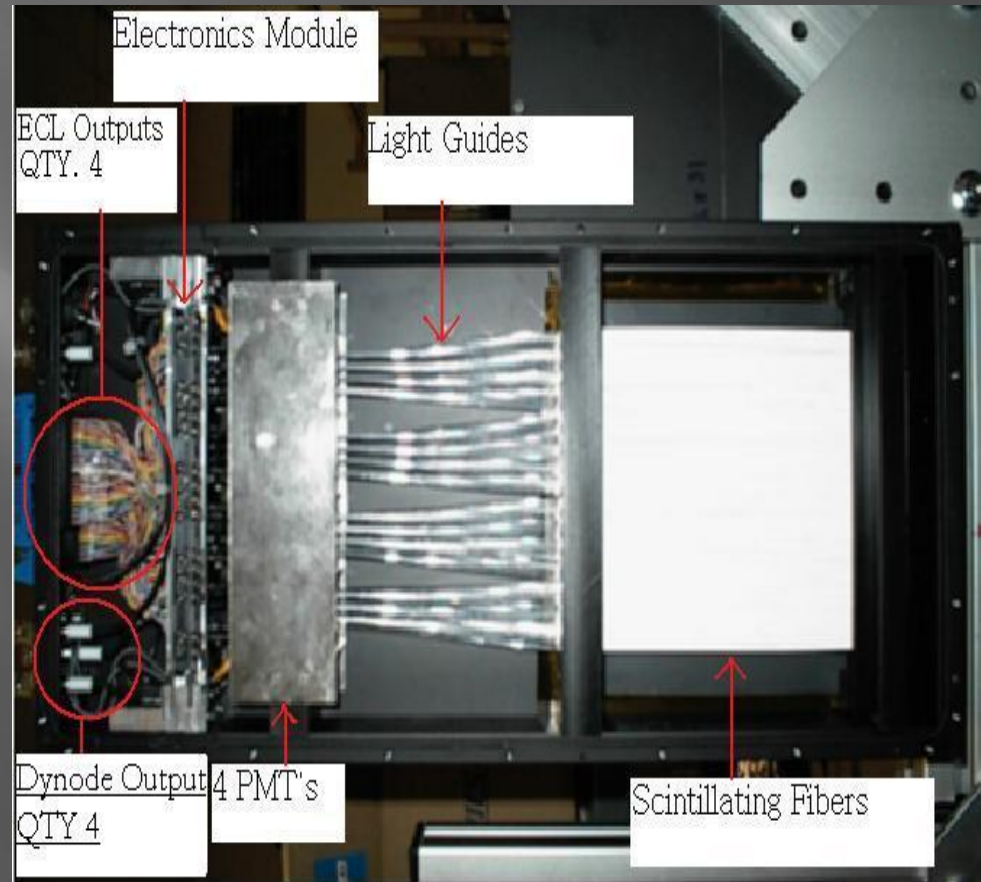
Testing Cont'd

- ❖ Purpose of the Beam Profile Monitor (BPM)
 - ❖ Determine the beam position and entrance angles as well as its' spatial and angular divergence
 - ❖ The detector is situated to measure the x and y coordinates of the incident photon beam (Horizontal and Vertical from the lab frame)



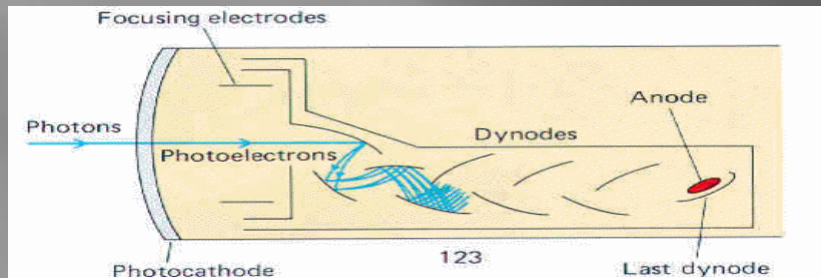
Testing Cont'd

- ❖ Internal structure of BPM (per module)
 - ❖ ~ 64 Scintillating Fibers (2mm x 2mm x 13 cm)
 - ❖ 4 groups of 16 light guides direct light into 4 photomultiplier tubes (PMT)
 - ❖ 16 anode signals per PMT are processed by a discriminator
 - ❖ One dynode signal per PMT



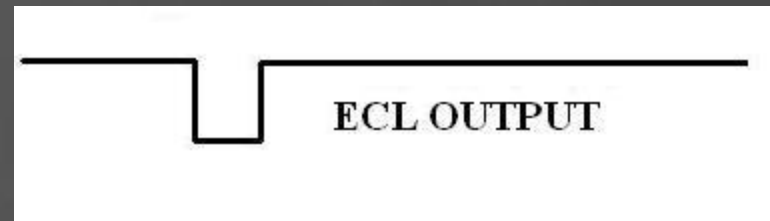
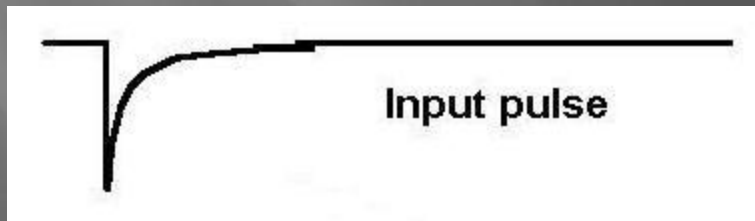
Testing Cont'd

❖ Photomultiplier Tube (PMT)



❖ Discriminator

- ❖ Only allows anode signals that meet the prescribed parameters



Testing Cont'd

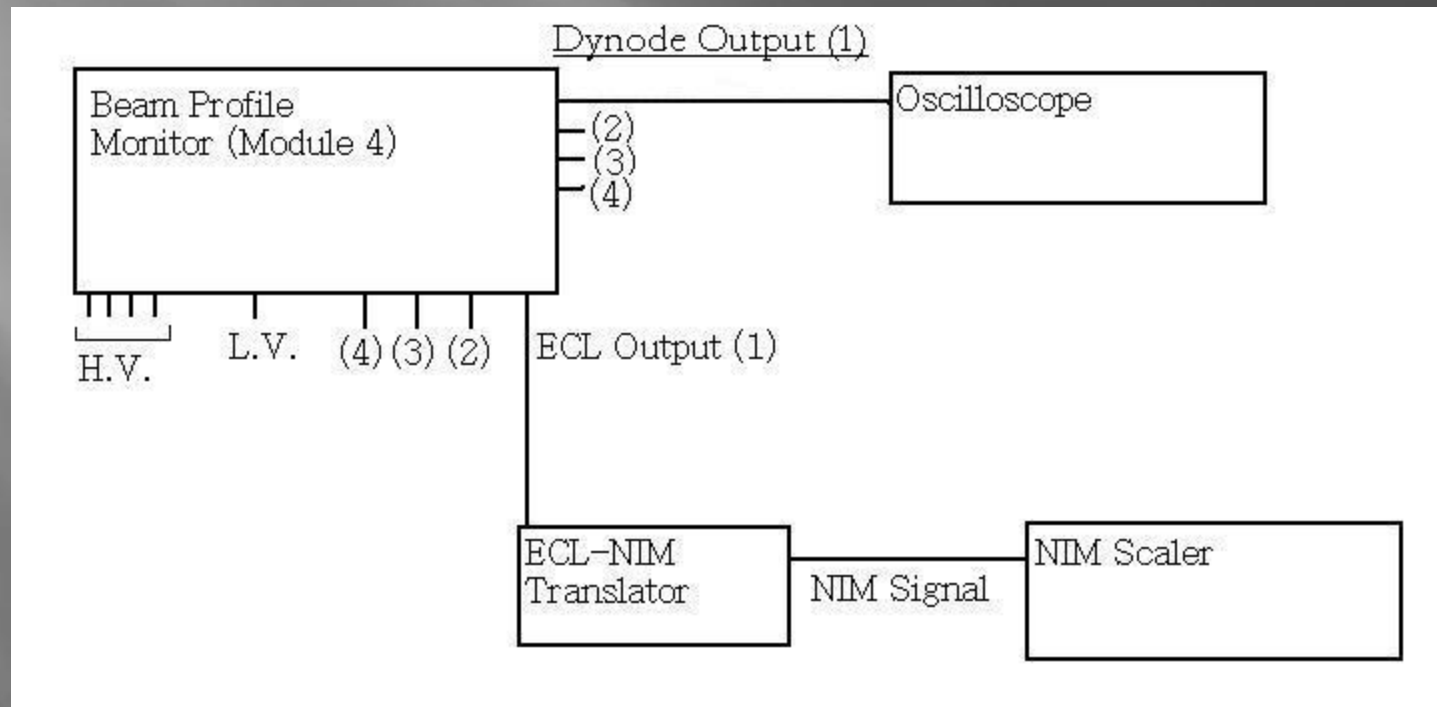
- ❖ ECL Output
 - ❖ Read on Scaler
 - ❖ Defined logical signal when event occurs
- ❖ Analog Output
 - ❖ Read on Oscilloscope
 - ❖ Shows a more detailed signal

Testing Cont'd

- ❖ Process
 - ❖ High voltage power supply ~ -850 V per PMT
 - ❖ Low voltage power supply (electronic module)
 - ❖ Connect detector to Oscilloscope and Scaler

Testing Cont'd

❖ Setup



Testing Cont'd

- ❖ Process Cont'd

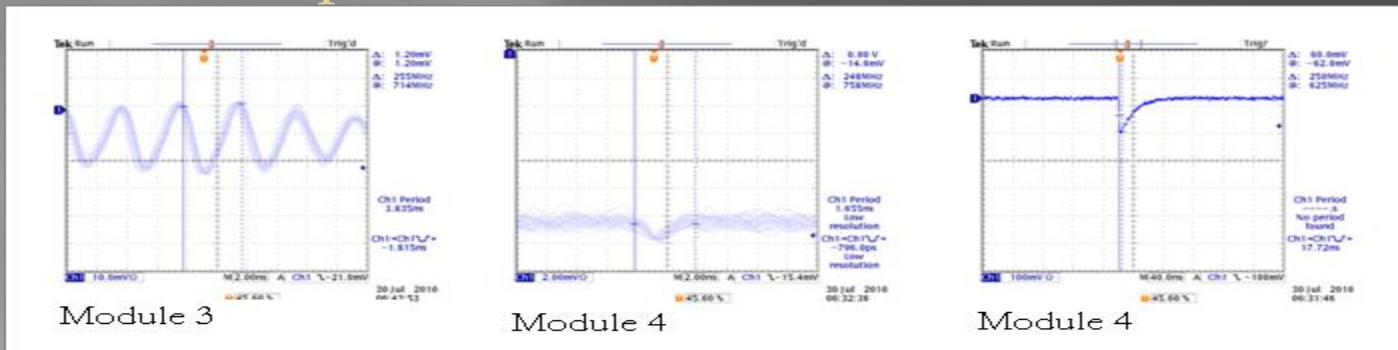
- ❖ Ensure detector is free from light leakage

- ❖ Check for faults

- ❖ Measure event rate on the scaler and event amplitude on oscilloscope

Testing Cont'd

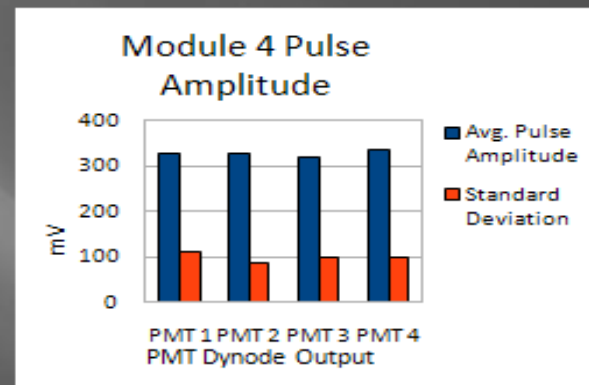
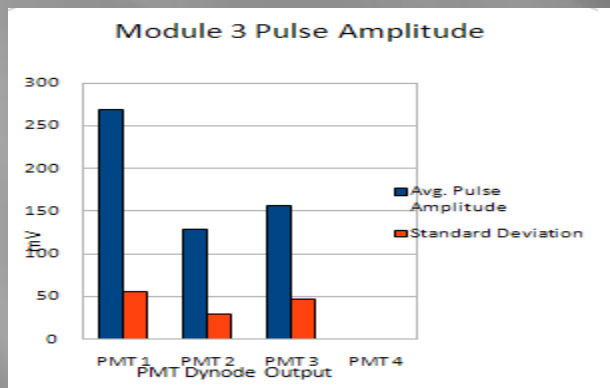
- ❖ Data Retrieved
- ❖ Oscilloscope



- ❖ Significant electronic noise on module 3
 - ❖ Thermionic emission, leakage current
- ❖ Significant after pulsing present
 - ❖ Residual gas, electrode glow
- ❖ No light leakage

Testing Cont'd

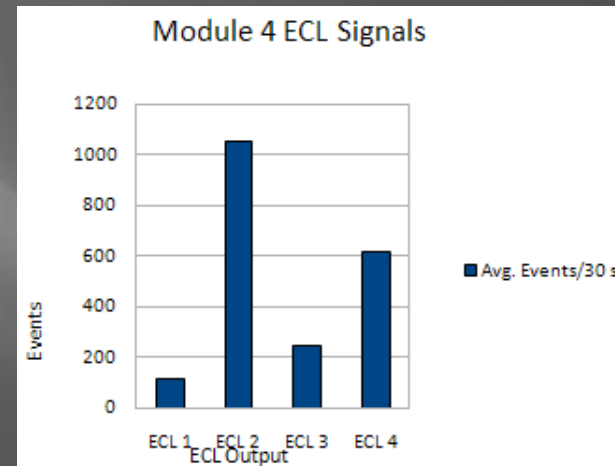
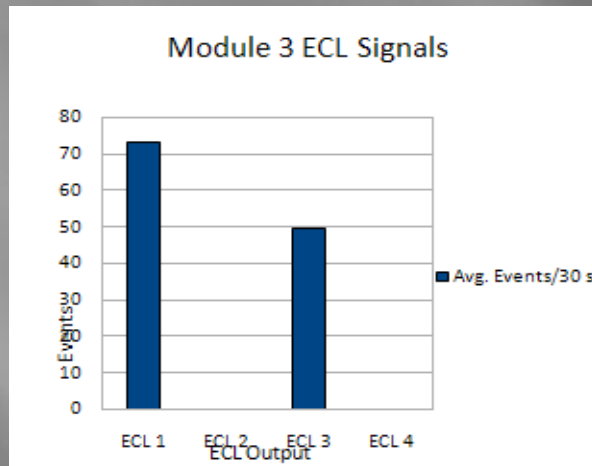
❖ Confirmation of results



- ❖ Strong detection and signal processing on all channels of module 4
- ❖ Gain deficiency on module 3 PMT's
- ❖ Further investigation required to determine exact cause of the problem

Testing Cont'd

❖ ECL Signals



❖ This data confirms the analog data

Testing Cont'd

- ❖ Summary of the test
 - ❖ PMT 4 on module 3 does not produce any signal
 - ❖ The discriminator for PMT 2 on module 3 requires adjustment
 - ❖ The high voltage power supply must be individually adjusted for each PMT on both modules
 - ❖ Each channel per PMT requires software calibration

Conclusion

- ❖ Muons are fundamental particles that interact via gravitation, electro-magnetism and the weak interaction
- ❖ Cosmic muons are produced when primary cosmic rays interact with earth's atmosphere
- ❖ Muon flux and energy is affected by atmospheric conditions
- ❖ Muons are beneficial to modern scientific research